

Name Example

Funniest Moment in Class _____

1. A star has a Luminosity of 4.5×10^{26} W, and a peak black body radiation of 520 nm. What is its radius? (8.08×10^8 m)

$$T = \frac{2.90 \times 10^{-3}}{\lambda}$$

$$T = \frac{2.90 \times 10^{-3}}{520 \times 10^{-9}} = 5576.92 \text{ K}$$

$$L = \sigma A T^4$$

$$4.5 \times 10^{26} = (5.67 \times 10^{-8}) 4\pi r^2 (5576.92)^4$$

$$r = 8.08017 \times 10^8 \text{ m}$$

2. A star has an apparent brightness of 2.35×10^{-14} W/m². If it has an absolute magnitude of 2.3, how many parsecs are we from it? (3590 → 3600 pc)

$$m = 2.5 \log \left(\frac{2.52 \times 10^{-8}}{2.35 \times 10^{-14}} \right) = 15.0758$$

$$m - M = 5 \log \left(\frac{d}{10} \right)$$

$$15.0758 - 2.3 = 5 \log \left(\frac{d}{10} \right)$$

$$10 \left(10^{\frac{m-M}{5}} \right) = 3590.59 \text{ pc}$$

3. A galaxy is 23 Mpc from us. At what wavelength would we see the 486 nm spectral line of Murrayium from that galaxy. (Use $H = 50$ km/s/Mpc) (487.863 → 488 nm)

$$v = Hd$$

$$v = \left(50 \frac{\text{km}}{\text{s}} \right) (23 \text{ Mpc}) = 1150 \frac{\text{km}}{\text{s}}$$

$$\frac{\Delta \lambda}{\lambda} = \frac{\Delta v}{c}$$

$$\frac{1150 \frac{\text{km}}{\text{s}}}{3 \times 10^8 \frac{\text{km}}{\text{s}}} = \frac{\Delta \lambda}{486}$$

$$\Delta \lambda = 1.863$$

$$\lambda' = 486 + 1.863$$

$$487.863 \text{ nm}$$

4. A very strong concertmaster is on top of a 148 m tall tower near a black hole where the gravitational field strength is 5.53×10^{12} m/s/s. If another player is making a frequency of 440.0 Hz at the bottom, what frequency does the concertmaster hear at the top? What beat frequency do they hear? (436. Hz, and 4.00 beats per second)

$$\frac{g \Delta h}{c^2} = \frac{\Delta f}{f}$$

$$\frac{(5.53 \times 10^{12})(148)}{(3 \times 10^8)^2} = \frac{\Delta f}{440}$$

$$\Delta f = 4.00 \text{ Hz}$$

Low clocks run slow so

the concert master hears $440 - 4 = 436 \text{ Hz}$

and this beats $f = |440 - 436| = 4.00 \text{ beats/sec}$

5. A black hole has a mass of 21 solar masses. (The mass of the sun is 1.99×10^{30} kg.) *Calculate the radius of the event horizon, *Calculate the time it would take a clock 160 km beyond the event horizon to register 60. seconds of elapsed time as we observe it from a great distance. (62 km, 71 s)

$$R_s = \frac{2GM}{c^2} = \frac{2(6.67 \times 10^{-11})(21)(1.99 \times 10^{30})}{(3 \times 10^8)^2} = 61942.1 \text{ m}$$

$$\Delta t = \frac{\Delta t_0}{\sqrt{1 - \frac{R_s}{r}}} = \frac{60 \text{ s}}{\sqrt{1 - \frac{61942.1}{(160,000 + 61942.1)}}} = 70.666 \text{ s}$$